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Galilean first-order formulation for the nonrelativistic expansion of general relativity

Dennis Hansen,1 Jelle Hartong,2 Niels A. Obers,3,4 and Gerben Oling3*
1Institut für Theoretische Physik, ETH Zürich, Wolfgang-Pauli-Strasse 27, 8093 Zürich, Switzerland
2School of Mathematics and Maxwell Institute for Mathematical Sciences, University of Edinburgh,
Peter Guthrie Tait Road, Edinburgh EH9 3FD, United Kingdom
3The Niels Bohr Institute, University of Copenhagen, Blegdamsvej 17, DK-2100 Copenhagen Ø, Denmark
4Nordita, KTH Royal Institute of Technology and Stockholm University,
Roslagstullsbacken 23, SE-106 91 Stockholm, Sweden

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We reformulate the Palatini action for general relativity in terms of moving frames that exhibit local Galilean covariance in a large speed of light expansion. For this, we express the action in terms of variables that are adapted to a Galilean subgroup of the GL(n, R) structure group of a general frame bundle. This leads to a novel Palatini-type formulation of general relativity that provides a natural starting point for a first-order nonrelativistic expansion. In doing so, we show how a comparison of Lorentzian and Newton-Cartan metric compatibility explains the appearance of torsion in the nonrelativistic expansion.

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I. INTRODUCTION

In recent years the study of nonrelativistic approximations to general relativity (GR) has gained renewed interest. As an important development, a geometric description of the nonrelativistic expansion of general relativity in inverse powers of the speed of light c was obtained in Refs. [1–3], building on earlier work [4,5]. To a large extent, this progress was made possible by new insights [6–8] in Newton-Cartan geometry [9,10], which takes the place of the Lorentzian geometry of GR at leading order in the nonrelativistic expansion. The resulting geometric nonrelativistic expansion of gravity can be carried out on the level of the action, which means that it can be applied to any astrophysical setting where velocities are small. In particular, the $1/c^2$ expansion of GR holds the prospect of being related to a covariant and off shell formulation of the post-Newtonian expansion, since the leading term has been shown to reproduce the 1PN approximation [5]. More generally, covariant theories of gravity involving Newton-Cartan-type geometries prominently appear in recent developments in nonrelativistic field theory [11–14], nonrelativistic string theory and limits of the AdS/CFT correspondence [15–17], along with many other areas that exhibit nonrelativistic physics.

In this work, we focus on the $1/c^2$ expansion of GR, which has been shown to lead to a modification of the original notion of Newton-Cartan (NC) geometry, known as type II torsional Newton-Cartan (TNC) geometry [2,3], as the correct framework for a covariant action of nonrelativistic gravity. When coupled to a point particle, the equations of motion of this action lead to the Poisson equation of Newtonian gravity in an arbitrary frame. In addition, for appropriate matter sources, the theory generalizes Newtonian gravity since it includes the effects of gravitational time dilation due to strong gravitational fields [1,3,18]. When time is absolute, type II TNC geometry reproduces the usual NC geometry.

To set the stage, we briefly review the nonrelativistic expansion of GR in powers of $1/c^2$ following Ref. [3]. We can make the factors of $c$ in the metric explicit by introducing the “prenonrelativistic” (PNR) parametrization

$$g_{\mu\nu} = -c^2 T_\mu T_\nu + \Pi_{\mu\nu}, \quad g^{\mu\nu} = -\frac{1}{c^2} V^\mu V^\nu + \Pi^{\mu\nu}. \quad (1)$$

The resulting PNR variables $(T_\mu, \Pi_{\mu\nu})$ and $(V^\mu, \Pi^{\mu\nu})$ are all of order $O(c^0)$. They satisfy the orthogonality and completeness relations

$$V^\mu T_\mu = -1, \quad T_\mu \Pi^{\mu\nu} = 0, \quad V^{\mu} \Pi_{\mu\nu} = 0, \quad (2a)$$

$$\delta^\mu = -V^\mu T_\mu + \Pi^{\mu\nu} \Pi_{\mu\nu}, \quad (2b)$$

which imply in particular that $\Pi^{\mu\nu}$ has rank $n - 1$, where $n$ is the spacetime dimension. Assuming analyticity, they can be expanded in powers of $1/c^2$, which leads to

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*Corresponding author.  
gerben.oling@nbi.ku.dk  

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$T_{\mu} = \tau_{\mu} + \frac{1}{c^2} m_{\mu} + O(1/c^4), \quad \Pi_{\mu\nu} = h_{\mu\nu} + \frac{1}{c^2} \Phi_{\mu\nu} + O(1/c^4)$. \hspace{1cm} (3a, 3b)

Likewise, $V^\mu = v^\mu + O(c^{-2})$ and $\Pi^{\mu\nu} = h^{\mu\nu} + O(c^{-2})$. The PNR variables transform under local Lorentz transformations, which reduce to Galilean transformations in the $1/c^2$ expansion. Under these Galilean transformations, $(\tau_{\mu}, h^{\mu\nu})$ are invariant, while $(v^\mu, h_{\mu\nu})$ transform.

Before expanding in powers of $1/c^2$, the prenonrelativistic variables simply provide a parametrization of the Lorentzian metric through Equation (1). Upon expanding, however, the resulting leading-order fields $\tau_{\mu}$ and $h^{\mu\nu}$ define a NC geometry, which is modified into a type II TNC geometry by adding the subleading fields $m_{\mu}$ and $\Phi_{\mu\nu}$. As such, the PNR parametrization (1) is a convenient way of recasting the Lorentzian metric variables of GR in such a way that the appropriate nonrelativistic geometry appears naturally in the expansion.

Next, we rewrite the Einstein-Hilbert action of GR,

$$S_{EH} = \frac{c^3}{16\pi G_N} \int_M \mathcal{L}_{EH} \sqrt{-g} d^4 x, \quad \mathcal{L}_{EH} = \mathcal{L}, \hspace{1cm} (4)$$

in terms of the PNR variables (1). Here, $R$ is the Ricci scalar of the Levi-Civita connection, which is the unique torsion-free connection that preserves the Lorentzian metric $g_{\mu\nu}$ under parallel transport, as is convenient for describing the degrees of freedom of GR. However, to obtain a connection that is compatible with the nonrelativistic geometry arising from the expansion (3), it is convenient to instead use the PNR connection

$$C^{\mu\nu}_A = -V^\rho \partial_\rho T_{\nu} + \Pi^{\rho\sigma} \partial_{(\rho} \Pi_{\sigma)\nu} - \frac{1}{2} \Pi^{\rho\sigma} \partial_\rho \Pi_{\sigma\nu}, \hspace{1cm} (5)$$

which leads to a covariant derivative under which $T_{\mu}$ and $\Pi^{\mu\nu}$ are covariantly constant. This has the consequence that the Newton-Cartan structure $(\tau_{\mu}, h^{\mu\nu})$ arising from these variables at leading order in the $1/c^2$ expansion is covariantly constant with respect to the covariant derivative coming from the leading-order term in the expansion of $C^{\mu\nu}_A$. Note that the connection (5) has nonzero torsion.

Up to a total derivative, the Einstein-Hilbert Lagrangian then takes the form

$$\mathcal{L}_{EH}^{(1)} = \frac{c^2}{4} \Pi^{\mu\nu} \Pi^{\rho\sigma} T_{\mu\nu} T_{\rho\sigma} + \Pi^{\mu\nu} R_{\mu\nu} - \frac{1}{c^2} V^\rho V^\sigma R_{\mu\nu}, \hspace{1cm} (6)$$

where $T_{\mu\nu} = (dT)_{\mu\nu} = \partial_\mu T_{\nu} - \partial_\nu T_{\mu} + \Ric_{\mu\nu}$ is the Ricci tensor corresponding to the connection (5). Note that $\sqrt{-g}$ is of order $c$, so the leading term in the action (4) is $O(c^0)$. The parametrization (6) is the starting point for the large speed of light expansion performed in Refs. [2,3]. So far, the resulting expansion of the action has been carried out up to second subleading order,

$$L_{EH}^{1} = c^2 L_{LO}^{1} + L_{NLO}^{1} + \frac{1}{c^2} L_{NNLO}^{1} + O(1/c^4). \hspace{1cm} (7)$$

The leading-order term in the expansion, coming from the first term in Eq. (6), imposes a particular constraint on the torsion of the geometry, as we will see below. The next-to-next-to-leading-order equations of motion (which include the equations of motion of the preceding orders [19]) describe a type II torsional Newton-Cartan theory of gravity. This theory reproduces Newtonian gravity and generalizes it to include strong gravity. However, carrying out the expansion to higher orders becomes increasingly cumbersome in a second-order formulation.

In this paper, our aim is to introduce the prenonrelativistic form of the first-order Palatini action of GR, which aids the computation of the $1/c^2$ expansion at higher orders. In addition, it is expected to simplify the study of a wide range of physical applications, for example by making the construction of boundary charges more accessible (see for example [20]), enabling the coupling to fermion fields, and simplifying Kaluza-Klein reductions.

Furthermore, we provide a new perspective on the geometric interpretation of the PNR parametrization using moving frames, which clarifies the appearance of torsion and the local Galilean symmetries associated to Newton-Cartan geometry. By embedding both the local Lorentz symmetry of GR and the local Galilean symmetry of NC geometry inside the GL(n, $\mathbb{R}$) frame bundle and its associated general (linear) affine connections, we can translate between the corresponding notions of torsion and metric-compatibility. This affine perspective also naturally connects to the “triality” between the usual metric formulation of GR and its equivalent formulations in terms of torsion or nonmetricity [21,22].

II. AFFINE AND LORENTZIAN CONNECTIONS

In the following, we will frequently use moving frames, which are also commonly known as vielbeine, to describe the geometry of the tangent bundle. For a spacetime $M$, we denote by

$$E^A = E^A_{\mu} dx^\mu, \quad \Theta_A = \Theta^\mu_A dx^\mu, \quad E^A(\Theta_B) = \delta_A^B, \hspace{1cm} (8)$$

a set of vielbeine and their duals, where $A = 0, \ldots, n - 1$ are tangent spacetime indices. Additionally, we introduce a connection $\Omega^A_{\mu\nu} = \Omega^A_{\mu\nu} dx^\mu$ and its associated covariant derivative $D$, which acts on frame tensors as

$$DX^A_B = dX^A_B + \Omega^A_{\mu\nu} X^C_B - \Omega^C_B X^A_C. \hspace{1cm} (9)$$

At this point, we take $\Omega^A_B$ to be a general (linear) affine connection, which takes values in the Lie algebra $\text{gl}(n, \mathbb{R})$. The torsion two-form $T^A$ of the connection is given by

$$T^A = dE^A + \Omega^A_B \wedge E^B. \hspace{1cm} (10)$$

Additionally, under local GL(n, $\mathbb{R}$) transformations $\Lambda^A_B$, the connection and a frame tensor $X^A_B$ transform as
Finally, the frame bundle connection $\Omega^A_B$ can be related to an affine connection $\Gamma^\mu_{\nu\rho}$ acting on tensor products of the tangent and cotangent bundle by requiring the vielbeine to be parallel with respect to the sum of these connections,

$$0 = \partial^\mu A^\nu - \Gamma^\mu_{\nu\rho} E^\rho_A + \Omega^A_B E^B_C,$$  \hspace{1cm} (13)

This relation is also known as the “vielbein postulate.”

So far, none of the above requires the existence of a metric. Introducing a Minkowski metric $\eta_{AB}$ on the frame bundle breaks the GL$(n, \mathbb{R})$ transformations of a general frame bundle down to the SO$(1, n-1)$ Lorentz transformations that leave it invariant. The Lorentz algebra also corresponds to the local symmetries of the associated Lorentzian tangent bundle metric

$$g_{\mu\nu} = \eta_{AB} E^A_\mu E^B_\nu.$$  \hspace{1cm} (14)

The metric $\eta_{AB}$ may not be covariantly constant for a general affine connection, which we can measure using

$$Q_{AB} = -D\eta_{AB}/2 = (\Omega_{AB} + \Omega_{BA})/2 = \Omega_{(AB)}.$$  \hspace{1cm} (15)

We therefore refer to $Q_{AB}$ as the Lorentzian nonmetricity tensor. Note that we have lowered one of the indices on the connection using $\eta_{AB}$, which allows us to consider its (anti-)symmetrization, so that we can split

$$\Omega_{AB} = \Omega_{(AB)} + Q_{AB}.$$  \hspace{1cm} (16)

$$T^A = dE^A + \Omega^{[AB]} \wedge E_B + Q^{AB} \wedge E_B.$$  \hspace{1cm} (17)

We can then use the torsion equation (17) to solve for $\Omega_{(AB)}$ in terms of the nonmetricity $Q_{AB}$ and torsion $T^A$,

$$\Omega_{AB} = \omega_{AB} + K_{AB} + L_{AB}.$$  \hspace{1cm} (18)

This expresses a general gl$(n, \mathbb{R})$ connection $\Omega^A_B$ in terms of the Levi-Civita connection, which we denote by

$$\omega_{AB} = dE_A(\Theta_B, \Theta_C)E_C - \frac{1}{2}dE_C(\Theta_A, \Theta_B)E_C,$$  \hspace{1cm} (19)

as well as the contorsion $K_{AB}$ and disformation $L_{AB}$,

$$K_{AB}(\Theta_C) = \frac{1}{2}T_C(\Theta_A, \Theta_B) - T_A(\Theta_B, \Theta_C),$$  \hspace{1cm} (20)

$$L_{AB}(\Theta_C) = \frac{1}{2}Q_{AB}(\Theta_C) - Q_{C(A}(\Theta_B)).$$  \hspace{1cm} (21)

Setting $T_A = 0$ and $Q_{AB} = 0$, the decomposition (18) implies the well-known fact that the Levi-Civita connection $\omega_{AB}$ is the unique torsion-free connection that preserves the Lorentzian metric $\eta_{AB}$.

Following Eq. (16), we have thus seen that introducing a frame metric such as $\eta_{AB}$ naturally splits an affine connection $\Omega^A_B$ into a part $\Omega_{(AB)}$ that preserves the metric and a part $Q_{AB} = \Omega_{(AB)}$ that does not. We will consider the Galilean version of this decomposition of a general gl$(n, \mathbb{R})$ connection below.

### III. PALATINI ACTION AND SHIFT

We now turn to general relativity, where our starting point is the frame formulation of the Palatini action,

$$S_{\text{Pal}}[E, \Omega] = \frac{1}{2\kappa c^2} \int_M \eta_{AB} \ast (E^A \wedge E^C) \wedge R^{B}_C,$$  \hspace{1cm} (22)

with $\kappa = 8\pi G_N c^{-4}$. This action contains both the vielbeine $E^A$ and a gl$(n, \mathbb{R})$ connection $\Omega^A_B$ as variables. The latter appears through the curvature two-form

$$R^A_B = d\Omega^A_B + \Omega^C_B \wedge \Omega^{CA}.$$  \hspace{1cm} (23)

Furthermore, the Hodge dual $\ast$ leads to a $(n - 2)$-form

$$\ast(E^A \wedge E^C) = \frac{\eta^{A_D} \eta^{C_E}}{(n - 2)!} e_D \wedge E^B \wedge \cdots \wedge E^D,$$  \hspace{1cm} (24)

so that the total integrand of the action (22) is an $n$-form. Here, $\epsilon_{A_1 \cdots A_n}$ is the fully antisymmetric symbol.

Although the Palatini action is formulated using the Minkowski metric $\eta_{AB}$, we actually do not need to assume that the connection variable $\Omega^A_B$ is metric compatible or torsionless. Starting from a fully general gl$(n, \mathbb{R})$ connection $\Omega^A_B$, its equations of motion lead to

$$\Omega_{AB} = \omega_{AB} + \eta_{AB} Z.$$  \hspace{1cm} (25)

The one-form $Z$ is arbitrary but drops out of the action upon substituting the solution, and we can remove this ambiguity using a Lagrange multiplier for a constraint, see for example [23,24]. Out of all possible gl$(n, \mathbb{R})$ connections, the Palatini action then uniquely selects the Levi-Civita connection (19). Substituting this solution in the action reproduces the Einstein-Hilbert action (4).

Our goal is now to identify a first-order formulation of the PNR parametrization of the Einstein-Hilbert action (6) in terms of the adapted connection $\Theta^A_B$ (5). This can easily be achieved using the linear field redefinition

$$\Omega^A_B = \tilde{\Omega}^A_B + S^A_B,$$  \hspace{1cm} (26)

where the “shift” parameter $S^A_B$ is a particular function of the vielbeine and their derivatives. The on shell value of the new connection variable $\tilde{\Omega}^A_B$ is then given by

$$\tilde{\Omega}^A_B|_{\text{on shell}} = \omega^A_B - S^A_B.$$  \hspace{1cm} (27)

As a result, even though the Palatini action (22) implies that the on shell value of $\Omega^A_B$ corresponds to the Levi-Civita connection, an alternative connection can be obtained on shell from $\tilde{\Omega}^A_B$ using a suitable shift $S^A_B$.

In the action (22), the field redefinition (26) leads to

$$S[E, \tilde{\Omega}] = \frac{1}{2\kappa c^2} \int_M \eta_{AB} \ast (E^A \wedge E^C) \wedge (\hat{R}^B_C + S^B_D \wedge S^D_C + \tilde{D}S^B_C).$$  \hspace{1cm} (28)
This “shifted” action now depends on $\tilde{\Omega}^A_{\ B}$ and its associated covariant derivative $\tilde{D}$. Through the field redefinition (26), it is equivalent to the Palatini action (22). In the following, we will identify a suitable shift $S^A_{\ B}$ to obtain the frame equivalent of the desired PNR connection (5). To show that the resulting action (28) is equivalent to the PNR parametrization (6) of the Einstein-Hilbert action, it is useful to rewrite it as

$$S[E,\tilde{\Omega}] = \frac{1}{2\kappa c} \int_M \Omega_{AB} \ast (E^A \wedge E^C) \wedge (\tilde{R}^B_{\ C} - S^B_{\ D} \wedge S^D_{\ C})$$

$$+ d[S^A_{\ B} \ast (E_A \wedge E^B)] + S^A_{\ B} \wedge D \ast (E_A \wedge E^B)$$

$$+ \ast (E^A \wedge E^C) \wedge (\Omega_{AB} + \Omega_{BA}) \wedge S^B_{\ C}. \quad (29)$$

Here, we have introduced a boundary term and reinstated the covariant derivative $D$ with respect to the original connection $\Omega^A_{\ B}$. On shell, the latter is the Levi-Civita connection, which implies $DE^A = 0$ and $Q_{AB} = 0$, so that integrating out the connection leads to

$$S[E] = \frac{1}{2\kappa c} \int_M \Omega_{AB} \ast (E^A \wedge E^C) \wedge [R^B_{\ C} - S^B_{\ D} \wedge S^D_{\ C}].$$

(30)

Up to a boundary term, this action is equivalent to the Einstein-Hilbert Lagrangian of GR, for any choice of the shift parameter $S^A_{\ B}$.

IV. GALILEAN CONNECTION AND ACTION

Just like how the Minkowski frame bundle metric $\eta_{AB}$ is associated to the Lorentzian tangent bundle metric $g_{\mu\nu}$ through Eq. (14), we can associate two $GL(n, \mathbb{R})$ tensors

$$t_A, \quad \pi^{AB}, \quad (31)$$

satisfying $t_A \pi^{AB} = 0$, to the PNR variables $T_\mu$ and $\Pi^{\mu\nu}$ defined in Eq. (1). Whereas the Minkowski metric $\eta_{AB}$ is fixed by the Lorentz subalgebra of $gl(n, \mathbb{R})$, the tensors (31) are left invariant by the Galilean algebra, which consists of spatial rotations and boosts with arbitrary velocity. We also introduce the dual tensors $t^A$ and $\pi_{AB}$ such that we have the orthogonality and completeness relations

$$t_A t^A = -1, \quad t^A \pi_{AB} = t_A \pi^{AB} = 0, \quad (32a)$$

$$\delta^A_B = -t^A t_B + \pi^{AC} \pi_{CB}, \quad (32b)$$

corresponding to the same relations for the metric variables in Eq. (2).

However, following the “prenonrelativistic” decomposition (1) of the Lorentzian metric $g_{\mu\nu} = \eta_{AB} E^A_{\mu} E^B_{\nu}$, we can decompose the Minkowski metric as

$$\eta_{AB} = -t_A t_B + \pi_{AB}, \quad \pi^{AB} = -t^A t^B + \pi^{AB}, \quad (33)$$

where we identify

$$t_A E^A_{\mu} = cT_\mu, \quad \pi_{AB} \theta^A_{\mu} \theta^B_{\nu} = \Pi_{\mu\nu}, \quad (34a)$$

$$t^A \theta^\mu_A = -V^\mu / c, \quad \pi_{AB} \theta^A_{\mu} \theta^B_{\nu} = \Pi^{\mu\nu}. \quad (34b)$$

The frame tensors $t_A$ and $\pi^{AB}$ (and hence also $T_\mu$ and $\Pi^{\mu\nu}$) can be shown to be invariant under local Galilean transformations, but their duals $t^A$ and $\pi_{AB}$ (and hence also $V^\mu$ and $\Pi_{\mu\nu}$) are not invariant.

We now want to recast the Palatini action (22) in terms of the PNR Galilean variables defined above. Before expanding in powers of $1/c^2$, the action and the frames are invariant under the Lorentz algebra, and this Galilean description is somewhat unnatural. However, as outlined in Sec. I, by rewriting the action for GR in this way, we are preparing ourselves for a covariant description of the nonrelativistic Newton-Cartan geometry and its Galilean symmetry that appears in the expansion.

Our main task now is to introduce a suitable Galilean connection, corresponding to the frame version of the $C^{\mu\nu}$ connection (5) that was mentioned in Sec. I, and to identify its associated shift $S^A_{\ B}$. The resulting connection is compatible with the Galilean metric structure (31) and has only the minimal amount of torsion required for a Galilean connection, as detailed below.

First, in analogy with the Lorentzian definition (15), we define the Galilean nonmetrics as

$$Q_A^{(G)} = -D t_A = t_B \Omega^B_{\ A}, \quad (35a)$$

$$Q^{BV}_{(G)} = D \pi^{AB}/2 = (\pi^{AC} \Omega^B_{\ C} + \pi^{CB} \Omega^A_{\ C})/2. \quad (35b)$$

It is useful to split the tangent space index $A = (0, a)$ into space and time components, so that we can write the Galilean tensors (31) as

$$t_A = \delta^0_A, \quad \pi^{AB} = \delta^0_A \delta^B_a \delta^{ab}, \quad (36)$$

$$E^A = (cT, E^a), \quad \Theta^A = (-V/c, \Theta^a). \quad (37)$$

The nonzero components of the nonmetrics (35) are

$$Q = \Omega^{0}_0, \quad Q_a = c^{-1} \Omega^0_a, \quad Q^{ab} = \Omega^{(ab)}. \quad (38)$$

Here and in the following, we have introduced the appropriate factors of $c$ so that the resulting objects such as $Q_a$ are $O(c^0)$. Any connection for which all these components vanish is compatible with the Galilean PNR structure, and will therefore lead to a Newton-Cartan metric-compatible connection upon expansion.

Note that we can raise and lower spatial indices using $\delta_{ab}$ and $\delta^{ab}$. Together with (38), this allows us to split a general $gl(n, \mathbb{R})$ connection into

$$\Omega^{[ab]}, \quad \Omega^{a}_0, \quad Q^{ab}, \quad Q, \quad Q_a. \quad (39)$$
in analogy with the Lorentzian decomposition (16). The torsion equations (10) then take the form
\[ T^0 = c dT + e Q_a \wedge \xi^a + c Q_T + T, \]
\[ T^a = d \xi^a + \Omega^{(ab)}_0 + \xi_b + c \Omega^a_0 \wedge T + Q^{(ab)}_0 + \xi_b. \] (40)

For Galilean connections with vanishing nonmetricity, the time component of the torsion \( T^0 = c dT \) is fixed independent of the remaining connection components [25]. This torsion is generically nonzero, and requiring that it vanishes would mean that the \( 1/c^2 \) expansion of the geometry does not accommodate time dilation.

Instead, we allow the time component of the torsion \( T^0 = c dT \) to be nonzero and only require that the spatial torsion \( T^a \) vanishes. As is known from a tangent bundle perspective [8], this combination of torsion constraints and Galilean metric compatibility does not fully fix the connection. Here, we choose to use the connection
\[ \Omega_{[ab]} = d \xi_{[a}(\theta_b, V) \xi_{[c]} - d \xi_{[a}(\theta_b, V) T - 1/2 d \xi_{c}(\theta_a, \theta_b) \xi^c, \] (41a)
\[ \Omega^a = c \Omega^a_0 = \frac{1}{2} d \xi^a(\theta_c, V) + d \xi^a(\theta_a, \theta_b) \xi^c. \] (41b)

This is the frame equivalent of the \( C_{ab} \) connection (5) that naturally appears in the prenonrelativistic decomposition of the Levi-Civita connection [3], as can be seen using the vielbein postulate (13). From an algebraic perspective, the connections (41) correspond to spatial rotations and Galilean boosts, see the Supplemental Material [27].

The Galilean connection (41) can be obtained as the on shell value (27) for \( \Omega^a \), using the shift
\[ S^a_b = \frac{c^2}{2} dT(\theta^a, \theta_b) T, \] (42a)
\[ S^0_a = \frac{c^2}{2} dT(\theta^a, \theta_c) \xi_c - c dT(\theta^a, V) T, \] (42b)
\[ S^0_b = \frac{c^2}{2} dT(\theta^a, \theta_c) \xi_c - c dT(\theta^a, V) T + \frac{1}{2c^2} [d \xi^a(\theta_b, V) + d \xi^a(\theta_b, \theta_c)] \xi^c. \] (42c)

We can now work out the terms in the second-order action (30) explicitly in terms of Galilean-covariant objects. First, we have
\[ - \eta_{ab} * (E^A \wedge E^C) \wedge S^D_a \wedge S^0_C \]
\[ = \epsilon_{abc} \epsilon_{def} \frac{c^2}{4} dT(\theta_a, \theta_c) dT(\theta_d, \theta_e), \] (43)
which is equivalent to the leading term in the Lagrangian (6).

The connection (41) is metric compatible, which means that all nonmetricities (38) vanish. As a result, the components of its curvature (23) are given by

\[ R^a_b = d \Omega^a_b + \Omega^a_c \wedge \Omega^b_c, \] (44a)
\[ R^a = c R^a_0 = d \Omega^a_b + \Omega^a_b \wedge \Omega^b. \] (44b)

The combination of these curvatures that enters in the action is invariant under local Lorentz transformations before expanding in powers of \( 1/c^2 \), and after expanding it is invariant under local Galilean boosts at each order.

We can now write the action with on shell connection (30) in a Galilean-covariant way as
\[ S = \frac{1}{2 \kappa} \int_M \left[ \frac{c^2}{4} dT(\theta_a, \theta_b) dT(\theta^a, \theta^b) + R^a_b(\theta_a, \theta^b) - \frac{1}{c^2} R^a_b(V, \theta_a) \right] E d^3 x, \] (45)
with \( E = \det(T_{\mu
u}, \xi^{(a)}_\mu) \). This action is still Lorentz invariant, but it is constructed using Galilean building blocks. It precisely reproduces the PNR form of the Einstein-Hilbert Lagrangian (6). Given the shift (42), the equivalent decomposition of the off shell action (29) in terms of a Galilean-compatible covariant derivative and nonmetricity can then be used for the nonrelativistic expansion of general relativity in the first-order formulation.

V. DISCUSSION AND OUTLOOK

Through an appropriate field redefinition, we have obtained an alternative formulation of the Palatini action, whose equations of motion give rise to a connection that is adapted to the Galilean-covariant Newton-Cartan structure that arises in a nonrelativistic \( 1/c^2 \) expansion of GR. By identifying the Lorentzian and Galilean nonmetricity associated to a general affine connection, we also see how our field redefinition naturally leads to nonzero torsion. After substituting the connection, the resulting action is equivalent to GR. As such, we can reproduce the geometric, off shell nonrelativistic expansion of GR that was recently developed in a second-order formulation [2,3], and our first-order approach will allow it to be extended to higher orders in a more tractable manner. Importantly, our first-order description is expected to be essential in establishing the connection between the \( 1/c^2 \) expansion of GR and the Post-Newtonian (PN) approximation (as shown to 1PN order in [5]). Furthermore, the expansion can also be applied to a wide range problems in astrophysics where velocities are low. We also expect that this action will be of use in computing conserved boundary charges, since this is typically more accessible in a first-order formulation.

Additionally, we can apply our procedure to the ultrarelativistic \( c \to 0 \) limit, where the Galilean local symmetry is replaced by the Carroll algebra [28,29]. The resulting expansion in powers of \( c^2 \) gives a novel approach to Carroll geometry (together with its subleading corrections), which is connected to the geometry of null surfaces in GR [30] and is therefore potentially relevant to the dynamics of...
black hole horizons [31] and the holographic description of asymptotically flat space at null infinity [32,33]. This geometry is constructed using the Carroll-invariant tensors $t^A$ and $\pi_{AB}$, which have an opposite index structure compared to their Galilei counterparts. In this sense, the Carrollian notion of metric compatibility is the dual of the Galilean case considered above, see also Ref. [34]. We will address the ultrarelativistic expansion of our novel action and the corresponding geometry in more detail in upcoming work [35]. Different approaches to frame and/or first-order formulations of nonrelativistic and ultrarelativistic limits and expansions of gravity have previously been considered in Refs. [29,36,37].

Finally, our field redefinition procedure can also be used to obtain Palatini-type actions for GR whose equation of motion lead to a connection with vanishing Riemann tensor but nontrivial torsion or nonmetricity. This corresponds to the teleparallel or symmetric teleparallel formulation of GR, respectively, see also [38].

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