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Torsional-rotational spectrum of doubly deuterated dimethyl ether (CH$_3$OCHD$_2$)

First ALMA detection in the interstellar medium

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ABSTRACT

Context. In 2013, we published the first rotational analysis and detection of mono-deuterated dimethyl ether in the solar-type protostar IRAS 16293-2422 with the IRAM 30 m telescope. Dimethyl ether is one of the most abundant complex organic molecules in star-forming regions, and its D-to-H (D/H) ratios are important to understand its chemistry and trace the source history.

Aims. We present the first analysis of doubly deuterated dimethyl ether (methoxy-d$_5$-methane, 1,1-dideuteromethylether) in its inter-vibrational state, based on an effective Hamiltonian for an asymmetric rotor molecule with internal rotors. The analysis covers the frequency range 0.15–1.5 THz.

Methods. The laboratory rotational spectrum of this species was measured between 150 and 1500 GHz with Lille’s submillimeter spectrometer. For the astronomical detection, we used the Atacama Large Millimeter/submillimeter Array observations from the Protostellar Interferometric Line Survey.

Results. New sets of spectroscopic parameters have been determined by a least squares fit with the ERHAM code for both symmetric and asymmetric conformers. As for the mono-deuterated species, these parameters have permitted the first identification in space of both conformers of a doubly deuterated dimethyl ether via detection near the B component of the Class 0 protostar IRAS 16293-2422.

Key words. molecular data – submillimeter: ISM – ISM: molecules – astrochemistry

1. Introduction

By May 2021, around 220 molecules have been detected in the interstellar medium (ISM) or circumstellar shells. Complex organic molecules (COMs) are ubiquitous in the ISM, especially close to the forming low-mass stars where the dust temperature is high enough to sublimate ice. With its nine atoms, dimethyl ether (CH$_3$—O—CH$_3$, hereafter DME), which was first detected in the ISM by Snyder et al. (1974), is a large COM of relevance for astrochemistry and one of the main COMs present in these warm and dense inner regions of the envelopes of Class 0 protostars, the so-called hot corinos (Ceccarelli et al. 2004). With the high sensitivity offered by new observations from the Atacama Large Millimeter/submillimeter Array (ALMA), new opportunities are opened up for characterizing the abundances of these COMs systematically (see, e.g., Jørgensen et al. 2020, for a recent review).

One of the characteristics of the molecules present in the warm gas in hot corinos is their high degree of deuteration. Although, the cosmic deuterium abundance relative to hydrogen is $1.5–2.0 \times 10^{-5}$ (e.g., Linsky 2003; Prodanović et al. 2010), the abundances of deuterated COMs relative to their non-deuterated counterparts are found to be much higher, up to $\approx 10\%$ or more in some cases (e.g., Parise et al. 2006; Jørgensen et al. 2018). These enhancements are thought to be a result of the exothermic reaction $\text{H}_3^+ + \text{H}_2 \rightleftharpoons \text{H}_2\text{D}^+ + \text{H}_2 + \Delta E$ that in gas at low temperatures enhances $\text{H}_2\text{D}^+$ relative to $\text{H}_3^+$. This effect is even more pronounced when one considers the multiple-deuterated variants, which for several species (e.g., D$_2$CO, Persson et al. 2018; CHD$_2$OHCHO, Manigand et al. 2019, D$_2$O, Jensen et al. 2021) are even further enhanced compared to the relative abundances of the mono- and non-deuterated species, possibly related to the buildup of the ice mantles where these species are formed.

Dimethyl ether is an interesting target species to extend these studies due to its structure and high abundance. The detection of the mono-deuterated form of CH$_3$OCH$_3$ toward the Class 0 protostellar system IRAS 16293-2422 was first reported by Richard et al. (2013) based on the detection of 20 lines of its symmetric and asymmetric conformers. Through ALMA observations, Jørgensen et al. (2018) determined the D/H ratio for these conformers (corrected for the number of equivalent hydrogen atoms)
2. Experimental details

2.1. Preparation of the doubly deuterated DME (methoxy-d₅-methane, 1,1-dideuteromethylether)

Potassium methoxide and p-toluensulfonyl chloride were purchased from Sigma-Aldrich. Methan-d₅-ol was purchased from Eurisotop.

Doubly deuterated dimethyl ether was synthesized by the reaction of potassium methoxide with 4-methylbenzenesulfonic acid, methyl-d₂ ester. The latter was prepared as previously reported for the trideutero derivative, but using methan-d₅-ol as alcohol (Yamamoto et al. 2016). In a three-necked flask under nitrogen and connected to a trap immersed in a cold bath of dry ice, potassium methoxide (5.6 g, 80 mmol) and dry DMSO (30 mL) were introduced. The mixture was heated to 60 °C and 4-methylbenzenesulfonic acid, methyl-d₂ ester (8.2 g, 40 mmol) in dry DMSO (20 mL) was added dropwise. The doubly deuterated DME was distilled off as it formed and condensed in the cold trap. At the end of the addition, the mixture was heated for 1 h at 80 °C. The trap was then closed with stopcocks. The yield was 1.0 g, thus 53%.

2.2. Nuclear magnetic resonance (NMR) of the doubly deuterated DME

1H NMR (CDCl₃, 400 MHz) δ 3.18 (quint, 1H, 2J_HD = 1.6 Hz, CHD₂); 3.23 (s, 3H, CH₃). 13C NMR (CDCl₃, 100 MHz) δ 59.5 (quint, 1J_C_D = 21.6 Hz, 1J_CH = 139.5 Hz (d), CHD₂); 60.1 (q, 1J_CH = 140.4 Hz, CH₃).

2.3. Lille THz spectrometer

The absorption spectra were measured between 150 and 1500 GHz using the Lille spectrometer (Zakharenko et al. 2015). The absorption cell was a stainless steel tube (6 cm in diameter and 220 cm in length). The measurements were performed at typical pressures of 35–60 Pa at room temperature. The different frequency ranges were covered with various active and passive frequency multipliers with the Agilent synthesizer (12.5–18.6 GHz) used as the primary signal source. To increase the sensitivity of the spectrometer, frequency modulation at 20.5 kHz of the reference source and lock-in detection were used. The demodulation of the detected signal may be performed either at 1f or 2f, but 2f demodulation is preferred because of the simpler presentation of the observed spectrum in this case. Absorption signals were detected by an InSb liquid He-cooled bolometer (QMC Instruments Ltd.). Estimated uncertainties for measured line frequencies are 30, 50, and 100 kHz depending on the observed signal-to-noise ratio (S/N) and the frequency range.

3. Spectral analysis

Similarly to the mono-deuterated species, the spectrum of the 2D-DME shows a very dense structure which is caused by the lines of the ground and the lowest excited vibrational states of symmetric and asymmetric conformations (See Fig 2). In addition, due to internal rotation of the unsubstituted methyl top, each rotational level of the two conformations is split into A and E symmetry components of the C₃v symmetry point group. Despite a relatively high barrier to internal rotation of V₁ ≈ 900 cm⁻¹ (see Table 1), the A – E splittings of rotational transitions were resolved for the majority of the ground state lines of the 2D-DME adding thus more complexity to the measured spectra. To treat the internal rotation, we used the ERHAM code (Groner 1997, 2012). The model used in the code refers to the so-called combined axes methods. It combines the rho-axis system in which the molecular Hamiltonian is set up, and the principal axes system to which the Hamiltonian is transformed. In the rho-axis system, the molecular internal z axis is set parallel to the ρ vector whose coordinates are calculated using the following expression:

\[ \rho_g = \frac{\lambda J_g}{I_g}, \quad (g = x, y, z). \]
where \( \lambda \) represents the direction cosines of the internal rotation axis of the top in the principal axis system, \( I \) represents the principal inertia moments, and \( \alpha \) is the inertia moment of the methyl top. The magnitude of the rho vector represents the coupling between internal rotation of the methyl top and overall rotation of the molecular frame. For the two conformations of 2D-DME, the rho values are relatively high (see Table 1). They represent an intermediate case of the torsion-rotation coupling in the line between cis methyl formate, HCOOCH\(_3\), \( \rho = 0.08 \) (Ilyushin et al. 2009), and acetaldehyde, CH\(_3\)C(O)H, \( \rho = 0.3 \) (Smirnov et al. 2014). For the latter two, strong torsion-rotation coupling, and medium height barriers to internal rotation significantly complicated the analysis and required the inclusion of many higher order terms into Hamiltonian. In the present analysis of the 2D-DME, a high barrier to internal rotation simplified the treatment, and in addition to pure rotational Hamiltonian only five terms were needed to describe a torsional-rotational interaction at close to experimental accuracy.

The symmetric conformation of 2D-DME has a symmetry plane and thus belongs to the \( C_\text{s} \) symmetry point group. Therefore, in the ERHAM code, the symmetry parameters ISCD is set to \(-1\) and \( \alpha \), and the angle between the rho vector and principal axes plane is set to 0. On the other hand, the asymmetric conformation has a minimal symmetry and belongs to the \( C_1 \) group. Consequently, for the asymmetric conformation, ISCD = +1 and after various attempts to set a value to \( \alpha \), we decided to fix the parameter to 0 for the same reasons as discussed in the Sect. 4 of the mono-deuterated analysis (Richard et al. 2013).

The analysis was started with a pattern recognition of a few \( K_a = 0 \) and \( K_a = 1^{+}K_{0,1} \) lines in order to produce a fit and a first prediction. Then, the fit was iteratively improved, releasing more parameters by adding new identified lines.

One difficulty encountered during the analysis was caused by several series of high \( K_\nu \) \( Q_{\nu-1} \) transitions. An example is given in Fig. 3 for the \( K_\nu = 24 \) series. In this case, the spectral lines in the band head represent a congestion of many rotational transitions with different \( J \) quantum numbers and different torsional symmetries. Such congestion has a strong influence on the spectral line shape. Taking the spectral congestion and the
second derivative spectrum into account, it leads to a significantly reduced line frequency measurement accuracy, typically three to ten times less accurate than usual. In the weighted least-squares fit implemented in the ERHAM code, the weight of each frequency is calculated to be reciprocal to the measurement accuracy squared. The influence of such highly congested lines would be ten to 100 times less important. Therefore, we decided to remove these congested lines from the final fit and to include only the lines from such a series that could be considered as isolated ones.

A total of 1451 lines with $J_{\text{max}} = 71$ and $K_a,\text{max} = 27$ and 2080 lines with $J_{\text{max}} = 72$ and $K_a,\text{max} = 25$ were assigned for symmetric and asymmetric conformers, respectively. For each conformer, a set of 15 pure rotational and five torsional-rotation parameters of the effective Hamiltonian implemented in the ERHAM code was needed to fit the observed molecular lines. The determined parameters are listed in Table 1. In the torsional-rotational part of the Hamiltonian, besides $\rho$, we determined the following parameters: $\beta$ the angle between rho vector and principal axis $a$, $\varepsilon_{10}$ the tunneling parameter representing the splitting between $A$ and $E$ symmetry sublevels, and two rotational constant tunneling parameters $[(A - (B + C)/2)\_\nu=1$ and $[(B - C)/4]_{\nu=1}$.

### Table 1. Spectroscopic constants of the ground-vibrational state of doubly deuterated DME for the two different conformers.

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<tr>
<th>Parameters</th>
<th>Symmetric conformer</th>
<th>Asymmetric conformer</th>
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<td>$A$ (MHz)</td>
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<td>34.199.90529(31)</td>
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<tr>
<td>$B$ (MHz)</td>
<td>9226.37143(19)</td>
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<td>$C$ (MHz)</td>
<td>8256.93333(16)</td>
<td>7984.12140(11)</td>
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<td>7.001791(72)</td>
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<td>-12.80820(49)</td>
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<td>$\Delta_2$ (kHz)</td>
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<td>208.1287(14)</td>
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<td>1.335158(25)</td>
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<td>0.0039288(15)</td>
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<tr>
<td>$\Phi_{1\nu}$ (Hz)</td>
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<td>0.0908(12)</td>
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<tr>
<td>$\Phi_2$ (Hz)</td>
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<td>$\Phi_3$ (Hz)</td>
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<td>6.8016(32)</td>
</tr>
<tr>
<td>$\phi_1$ (Hz)</td>
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<td>0.001907(62)</td>
</tr>
<tr>
<td>$\phi_{1\nu}$ (Hz)</td>
<td>1.779(83)</td>
<td>2.857(41)</td>
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<tr>
<td>$\phi_3$ (kHz)</td>
<td>167.2067(20)</td>
<td>208.1287(14)</td>
</tr>
</tbody>
</table>

### Notes. Numbers in parentheses are one standard deviation in the same units as the last digit. (a) By symmetry. (b) Assumed, see Sect. 4 in Richard et al. (2013). (c) Barrier height for rotation of the methyl group determined with the program BARRIER. See Sect. 4 for uncertainties regarding the computation. (d) Number of distinct lines in the fit. (e) Error of the poorest fit line. (f) Number of free parameters used in the fit. (g) Standard deviation of the fit. (h) Weighted standard deviation of the fit.

### 4. Discussion

This study represents the first comprehensive characterization of the rotational spectrum of the two conformations of 2D-DME (CH$_2$OCHD$_2$) in the ground vibrational state, and in the frequency range up to 1.5 THz. Experimental measurements, partially given in the appendices in Tables A.1 and A.2, are available in their entirety in electronic form at the Centre de Données astronomiques de Strasbourg (CDS).

In the ERHAM code, the barrier to internal rotation is not calculated explicitly. Consequently, we used the program BARRIER which is derived from the program ASTOR described in Groner et al. (1986) that determines the potential barrier for a single simple internal rotor from torsional transitions or splittings. The ground state energy splitting between the $A$ and $E$ symmetry torsional substates, $\Delta E(E - A)$, was combined with the derived value of the internal rotation constant $F$ to obtain an estimate of the potential barrier $V_3$ for both conformers. These values are the direct output of ERHAM. In particular, the internal rotation constants, $F$, is obtained by ERHAM from $\rho$, $\beta$, and the rotational constants. The $F$ values determined in this study are 6.223 and 6.338 cm$^{-1}$ for symmetric and asymmetric conformers, respectively. The determined $V_3$ values are given in Table 1. The estimated error of the $V_3$ barriers was derived with the assumption that the relative error of $V_3$ is the same as the relative error of the $\Delta E(E - A)$ energy difference and the tunneling coefficient $\varepsilon_{10}$. We found an uncertainty of 3 and 1.5 cm$^{-1}$ for the symmetric and asymmetric conformers, respectively. The determined barrier heights agree well with the results of the study of monodeuterated DME (Richard et al. 2013), and with earlier studies of parent DME (Durig et al. 1976; Lovas et al. 1979).

Under Doppler limited spectral resolution, we did not observe any additional splittings in the ground state rotational lines of the asymmetric conformation as it could be expected, owing to possible tunneling between its two equivalent configurations, and as it was previously observed for monodeuterated DME (Richard et al. 2013). It should be noted that for the latter, the tunneling splittings were barely resolved for a limited number of lines. The lack of additional tunneling splittings for the asymmetric conformer may be explained by the increased mass of the internal rotor which is composed of two deuteriums and one hydrogen for 2D-DME. Provided that the barrier height does not vary significantly under H to D isotopic substitution, the increased internal rotor mass reduces the tunneling probability and consequently the splitting between the two tunneling substates (compared to monodeuterated DME). In this case, one could also expect smaller and unresolvable splittings under the Doppler limit in the rotational spectra. The absence of resolved additional splittings as well as the weighted rms deviations of the fits close to 1 support our initial strategy to treat the two conformations independently as two asymmetric rotors each having a single methyl top.
As an application of the new spectroscopy, we searched for the deuterated species toward CH$_3$OCHD$_2$, here, we targeted a position offset by 0.5" to the southwest from the "B" component of the system (IRAS16293B) – also the target of a number of the previous PILS papers. The advantage of searching at that position is that the lines are narrow ($\approx 1$ km s$^{-1}$) and the opacity of both the main lines of the dust continuum is limited. This enables easy assignments of individual transitions without problems with line overlaps and absorption.

At the targeted scales, the densities should be high enough for the excitation of the molecules to occur under local thermodynamic equilibrium (LTE) conditions. We consequently used synthetic spectra calculated under this assumption to identify the lines of CH$_3$OCHD$_2$ within the PILS range and used them to constrain their column density. The synthetic spectrum method is preferred over classical rotation diagrams, for example, as it immediately makes it possible both to search for good matches in the data and to show if lines are predicted at frequencies where no emission is seen. In particular, confirming the presence of rare species to demonstrate that the latter is not the case is equally important in finding good assigned matches. Also, this method makes it possible to predict the profiles for partially blended lines and also to look for overlaps with previously identified species.

Under the LTE assumption, the line profiles are calculated assuming Gaussian line profiles with constant LSR velocity offsets and line widths. Also, the extent of the source emission (and consequently beam filling factor) is kept constant. For these parameters, we adopted the same values as for the single-deuterated variants and most other species in the PILS range – a LSR velocity of 2.6 km s$^{-1}$, a line width (FWHM) of 1 km s$^{-1}$, and source size of 0.5" (assumed as a Gaussian distribution). The choices of the LSR velocity and line width can be tested directly when comparing to the observed spectra. The assumed source size mainly affects the results when optically thick lines are directly when comparing to the observed spectra. The assumed source size mainly affects the results when optically thick lines are immediately made it possible both to search for good matches in the data and to show if lines are predicted at frequencies where no emission is seen. In particular, confirming the presence of rare species to demonstrate that the latter is not the case is equally important in finding good assigned matches. Also, this method makes it possible to predict the profiles for partially blended lines and also to look for overlaps with previously identified species.

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Fig. 4. Spectra for the 24 transitions of the symmetric conformer of CH$_3$OCHD$_2$ predicted to be the strongest given the derived excitation temperature and column density. In the panels, the red line shows the fit to the CH$_3$OCHD$_2$ transition while the blue lines show the predictions for of all other species identified as part of the PILS survey (see, e.g., Drozdovskaya et al. 2019; Manigand et al. 2020, and references therein).
temperature. For the latter, we adopted a temperature of 125 K of the non- and singly-deuterated variants of dimethylether, again under the assumption that the doubly deuterated variants are present in the same gas. We then performed χ²-fitting to the reasonably isolated lines to constrain the column densities of the symmetric and asymmetric conformers separately.

Figures 4 and 5 show the 24 lines predicted in this manner to be the strongest for the asymmetric and symmetric conformers. For the two species, good fits can be seen for about 20 transitions – a number of which are well-isolated with peak intensities of up to about 200 mJy beam⁻¹. The derived column densities are $6.1 \times 10^{15}$ cm⁻² for the asymmetric conformers.

Fig. 5. Spectra for the 24 transitions of the asymmetric conformer of CH₃OCHD₂ predicted to be the strongest given the derived excitation temperature and column density.
and \(2.7 \times 10^{15} \text{ cm}^{-2}\) for the symmetric conformers. A priori, one would expect a factor 2 difference between the two due to the symmetry that occurs when two deuterium atoms replace two of the three hydrogen atoms in a CH$_3$-group of dimethyl ether. The factor 2.3 we derived when fitting the two conformers independently is consequently an additional verification of the assignment of the two conformers.

Taken together, the derived column densities imply a ratio between the column densities of the doubly to mono deuterated forms, \(N(\text{CH}_3\text{OCHD}_2)/N(\text{CH}_3\text{OCH}_2\text{D})\) of 15–20% for the two conformers. These imply D/H ratios that are about a factor of 5 higher than those inferred from the singly to non deuterated variants once corrected for the number of equivalent hydrogen atoms (e.g., Appendix B of Manigand et al. 2019). The trend of strong enhancements of the multi-deuterated variants also seen for other COMs thus also holds for DME, and they provide additional constraints on their formation during the cold phases of star formation.

7. Conclusion

Here, we have presented a complete high-resolution study of the doubly deuterated dimethylether over a decade in the frequency range 150–1500 GHz (0.15–1.5 THz) in its two conformations, symmetric and asymmetric. We were able to determine an accurate set of spectroscopic parameters that allowed us to reproduce measurements with a standard deviation better than 80 kHz. Thanks to these frequency predictions, both conformers have been detected near the “B” component of the protostellar system IRAS 16293-2422. The doubly deuterated species is found to be enhanced above the D/H ratio inferred from the mono-to-singly deuterated variants at a level similar to other organic molecules. These enhancements make CH$_3$OCH$_3$ an interesting target for further spectroscopic studies of other multiple deuterated variants, which in turn may shed light on the formation of this complex organic during the cold pre- and protostellar stages.

Acknowledgements. This work has been supported by the “Programme National Physique et Chimie du Milieu Interstellaire” (PCMI) of CNRS/INSU with INCINP co-funded by CEA and CNES. This project has received financial support from the CNRS through the MTTI interdisciplinary programs. J.C.G. thanks the Centre National d’Etudes Spatiales (CNES) for a grant. The research of J.K.J is supported by the Independent Research Fund Denmark (grant number 0135-00123B).

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### Appendix A: Experimental frequencies

#### Table A.1. Experimental frequencies measured in laboratory up to 1.5 THz for the symmetric conformer.

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<tr>
<th>σ</th>
<th>J′</th>
<th>K′</th>
<th>K′′</th>
<th>J′′</th>
<th>K′′</th>
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<th>Uncertainty (MHz)</th>
<th>o.-c. (MHz)</th>
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**Notes.** (a)Symmetry number: 0(A), 1(E).

#### Table A.2. Experimental frequencies measured in laboratory up to 1.5 THz for the asymmetric conformer.

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**Notes.** (a)Symmetry number: 0(A), 1(E).
## Appendix B: Predicted frequencies

### Table B.1. Predicted transition frequencies of doubly deuterated DME in the ground-vibrational state for the symmetric conformer.

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**Notes.** \(^{(a)}\) Symmetry number: 0(A), 1(E).

### Table B.2. Predicted transition frequencies of doubly deuterated DME in the ground-vibrational state for the asymmetric conformer.

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**Notes.** \(^{(a)}\) Symmetry number: 0(A), 1(E).